

0. Introduction – 0.1 Concept

The air / environment (geosphere): Is it a reactor ?

It's a matter of reactions and transports and mixing !

mixing times:

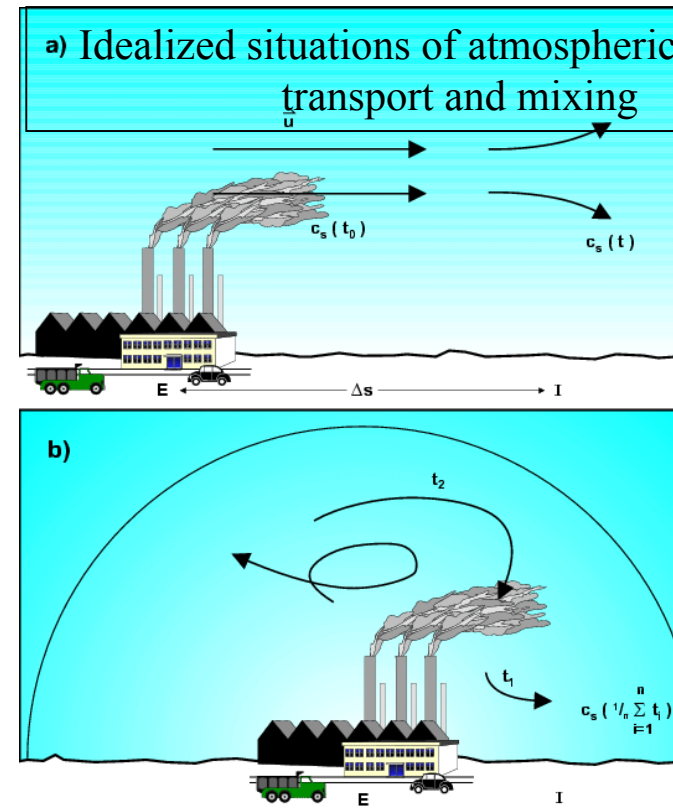
- vertically lower few kilometers (boundary layer) 1h-1d, mixing with free troposphere 2-10 days
- around the globe on the same latitude (zonal transport) within a few weeks
- from mid latitudes to the pole (meridional transport) within days to weeks, hemispheric mixing 2-6 months
- Between hemispheres about 1 year
- troposphere-stratosphere 1-3 years

Conceptually / knowledge to understand:

- (chemical) reactions
- (meteorological) transports and mixing

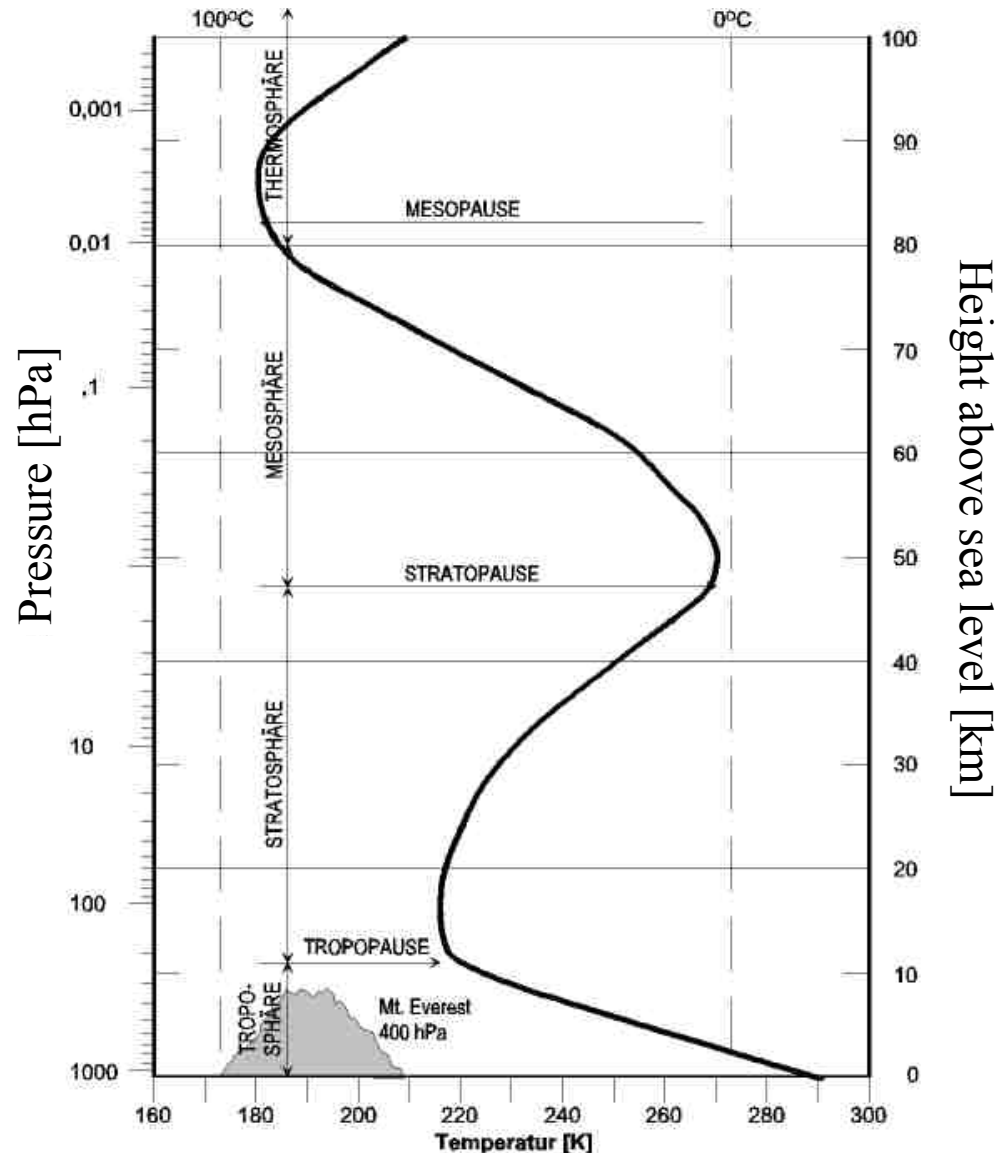
Tools to understand:

- 3D transport models
- model parameters isolated in laboratory experiments
- significant ,ingredients‘ identified in the ,field‘



0.2 Atmospheric pressure and composition

Pressure and temperature profiles in the atmosphere



Pressure decline

$$\Delta p = -g \rho \Delta z$$

$$\rho = 1.225 \text{ mg/cm}^3$$

$$\rightarrow \Delta z / \Delta p = 8 \text{ m/hPa}$$

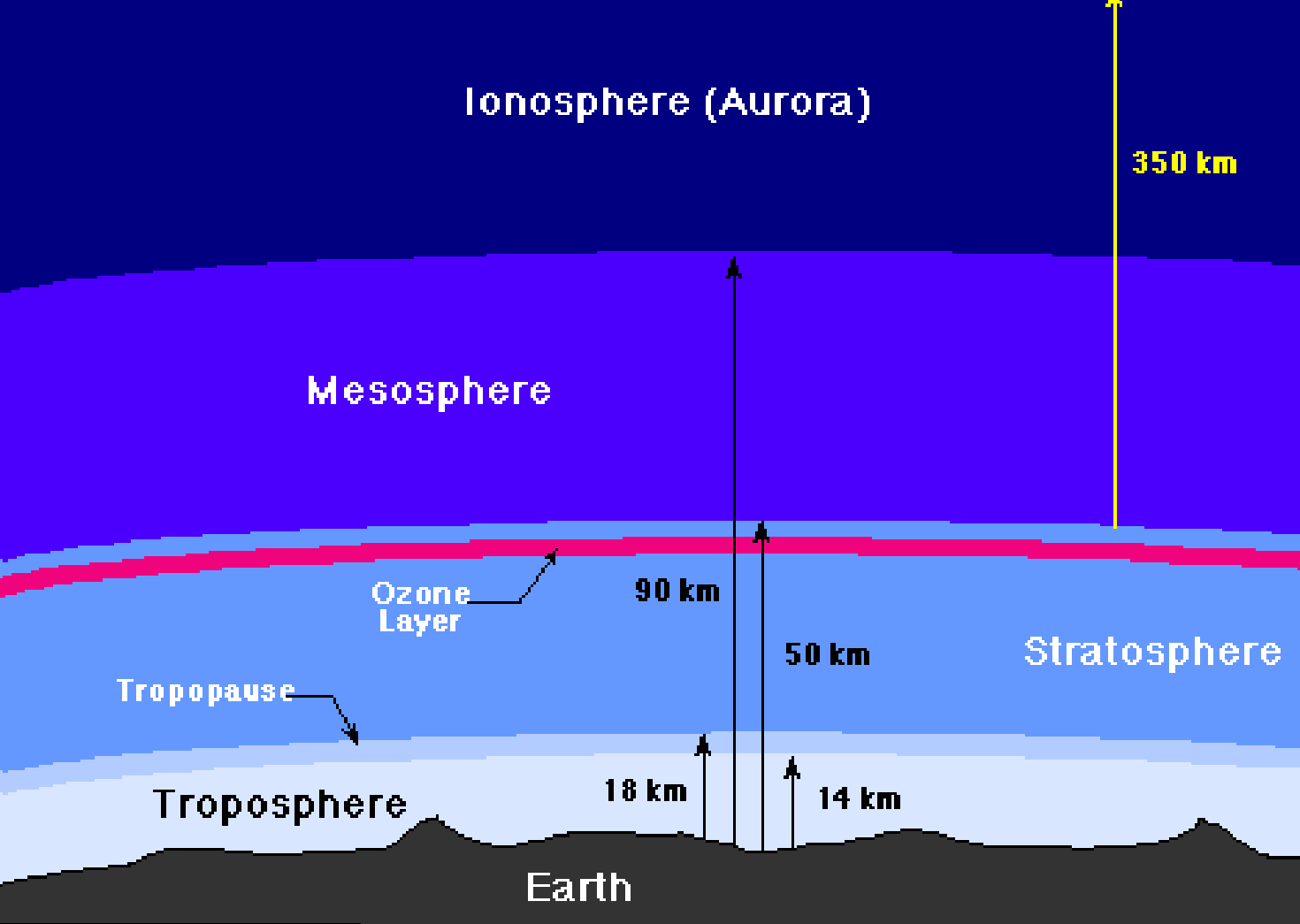
$$16 \text{ m/hPa in 6k}$$

„barometric st

$$p(h) = p_0 e^{-hg/RT}$$

, Standard

km	T [°C]
0	0
0	15
0	25
3	-4.4
10	-49.9
20	-56.5
30	-46.6
100	-63



3	-	4.4	700	28.964	1.76×10^{19}
10	-	49.9	265	28.964	6.67×10^{18}
20	-	56.5	55	28.964	1.38×10^{18}
30	-	46.6	12	28.964	0.30×10^{18}
100	-	63	.00021	28.5	5.29×10^{12}

0.3 Units for quantification of atmospheric trace substances

Ideal gas law: $pV = nRT = mRT/M_g$

Universal gas constant $R = 8206 \text{ Pa L}/(\text{mol K}) = 8.314 \text{ J}/(\text{mol K}) = N_A k_B$

Avogadro's number $N_A = 6.023 \times 10^{23}$

Boltzmann constant $k_B = 1.38 \times 10^{-23} \text{ J/K}$

$$1 \text{ Pa} = 1.013 \text{ N/m}^2 = 1.013 \text{ J/m}^3$$

Molar mass $M_{g \text{ air}} \approx 28.8 \text{ g/mol}$

Molar volume at $T_0 = 273 \text{ K}$ and $p_0 = 101325 \text{ Pa}$: $V = 22.414 \text{ L/mol}$

Concentration $c_i = m_i/V [\mu\text{g/m}^3]$ (for gases: = density)

Mass mixing ratio $\mu_{m i} = c_i/c [\text{ , \%}, \text{ppmm}, \text{ppbm}]$

Partial pressure $p_i = n_i RT/V_i [\text{Pa}]$

Volume mixing ratio $\mu_{V i} = p_i/p = V_i/V [\text{ , \%}, \text{ppmV}, \text{ppbV}]$

Number density $N_i/V = n_i N_A/V = p_i N_A/RT [\text{molec/cm}^3]$

Avogadro's number $N_A = 6.022 \times 10^{23} \text{ molec/mol}$

More details: Schwartz & Warneck (1995): Units for use in atmospheric chemistry, Pure Appl. Chem. 67, 1377-1406

Chemical Composition of the Atmosphere

<i>Constituent</i>	<i>Chemical Formula (sum)</i>	<i>Volume Mixing Ratio in Dry Air 10⁻¹³ 1!</i>	<i>Major Sources and Remarks</i>
Nitrogen	N ₂	78.084%	Biological
Oxygen	O ₂	20.948%	Biological
Argon	Ar	0.934%	Inert
Carbon dioxide	CO ₂	360 ppmv	Combustion, ocean, biosphere
Neon	Ne	18.18 ppmv	Inert
Helium	He	5.24 ppmv	Inert
Methane	CH ₄	1.7 ppmv	Biogenic and anthropogenic
Hydrogen	H ₂	0.55 ppmv	Biogenic, anthropogenic, and photochemical
Nitrous oxide	N ₂ O	0.31 ppmv	Biogenic and anthropogenic
Carbon monoxide	CO	50-200 ppbv	Photochemical and anthropogenic
Ozone (troposphere)	O ₃	10-500 ppbv	Photochemical
Ozone (stratosphere)	O ₃	0.5-10 ppm	Photochemical
Nonmethane hydrocarbons		5-20 ppbv	Biogenic and anthropogenic
Halocarbons (as chlorine)		3.8 ppbv	85% anthropogenic
Nitrogen species	NO _y	10 ppt-1 ppm	Soils, lightning, anthropogenic
Ammonia	NH ₃	10 ppt-1 ppb	Biogenic
Particulate nitrate	NO ₃ ⁻	1 ppt-10 ppb	Photochemical, anthropogenic
Particulate ammonium	NH ₄ ⁺	10 ppt-10 ppb	Photochemical, anthropogenic
Hydroxyl	OH	0.1 ppt-10 ppt	Photochemical
Peroxyl	HO ₂	0.1 ppt-10 ppt	Photochemical
Hydrogen peroxide	H ₂ O ₂	0.1 ppb-10ppb	Photochemical
Formaldehyde	CH ₂ O	0.1-1 ppb	Photochemical
Sulfur dioxide	SO ₂	10 ppt-1 ppb	Photochemical, volcanic.

1 Reaction types, kinetics

1.1 Reaction rate coefficient k (dt.: Reaktionsgeschwindigkeitskoeffizient)

Temperature dependence:

$$-dc_i/dt = k_T c_i$$

$$k_T = A(T) e^{[\Delta E/(RT)]}$$

Universal gas constant $R = k_B N_A = 1.38 \times 10^{-23} \text{ J/K} \times 6.023 \times 10^{23} / \text{mol}$

Providing the energy is sufficiently large, the temperature dependence of A

is unimportant, and k_T follows the Arrhenius expression:

$$k_T = A e^{[E_a/(RT)]}$$

with activation energy E_a

Frequently used, too: van t'Hoff expression:

$$k_T = B e^{[-E_a/R (1/T - 1/T_{\text{ref}})]},$$

The two expressions are equal via: $A = B e^{[E_a/(RT_{\text{ref}})]}$

Rate coefficient from gas kinetic theory

= effective reactions per unit of time =

= collision frequency x likelihood of reaction per collision

1) Molecules need to collide ($\approx 10^9$ /s at 10^5 Pa).

They are close enough to collide during $\approx (10^{-13}$ - $10^{-12}) \times 10^9$ /s

Collision frequency

$$Z_{ij} = \left(\frac{8kT}{\pi m_{ij}} \right)^{1/2} \sigma_{ij}^2 N_i N_j \quad \text{cm}^{-3} \text{s}^{-1}$$

$$\left(\frac{8kT}{\pi m_{ij}} \right)^{1/2} \quad \text{root - mean - square relative speed of molecules } i \text{ and } j$$

$$m_{ij} = \frac{m_i m_j}{m_i + m_j} \quad \text{reduced mass}$$

N_i : number density (cm^{-3})

$k = 1.381 \cdot 10^{-23} \text{ JK}^{-1}$ Boltzmann constant

σ_{ij}^2 : collision cross section

Speed, collisions in absolute terms:

from Maxwell-Boltzmann distribution function of molecular velocities (gas kinetic theory):

$$dN/N = c^{(m \langle v_i \rangle^2 / 2kT)}; \text{ with } c = (m / 2\pi kT)^{0.5}$$

$\langle v_i \rangle$ = mean absolute velocity of molecules

corresponding mean free path λ with regard to collisions

$$\lambda_i = 1 / [(N_A/V)(\sqrt{2} \pi \sigma_i^2)],$$

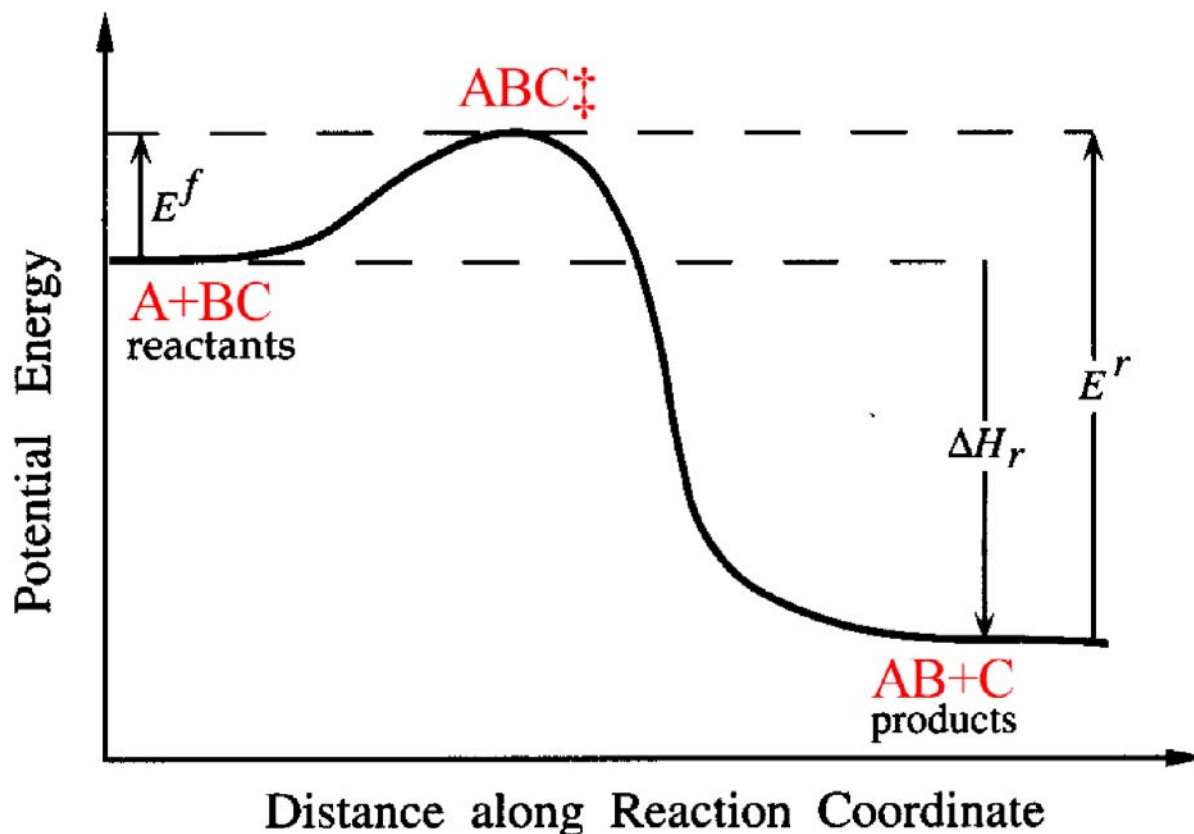
$$N_A/V = 6.022 \times 10^{23} / 0.0224 = 2.46 \times 10^{25} \text{ molec/m}^3$$

$$N_2 : \langle v \rangle = 0.51 \text{ m/s} = 1840 \text{ km/h}, \lambda = 65 \text{ nm}$$

$$CO_2: \langle v \rangle = 0.41 \text{ m/s} = 1490 \text{ km/h}, \lambda = 44 \text{ nm}$$

2) Extra energy is needed to overcome the transition state

As molecules approach each other they have to overcome an energy barrier because of electronic repulsion. The energy they absorb leads to a loosening of the intramolecular bonds until a transition state is reached, which can then either fall back into the original molecules or proceed to yield the products of the reaction. Reaction enthalpy ΔH_r .



Likelihood: collision cross section, σ_{ij}

In the transition state, usually the energy provided from bond formation is less than the energy consumption from bond breaking.

The collision cross section takes into account that a minimum amount of energy is needed. It is not necessarily provided with every orientation of the molecules during approach (steric factor).

$$\sigma_{ij} = \pi(r_i + r_j)^2 \left(1 - \frac{E_0}{E}\right) \quad \text{for } E \geq E_0, \quad 0 \text{ elsewhere}$$

Integrated over all total energies from 0 to ∞ , the frequency of reactive collisions becomes

$$Z_{ij}^R = \pi(r_i + r_j)^2 \left(\frac{8kT}{\pi m_{ij}}\right)^{1/2} e^{-E_0/kT} N_i N_j$$

k

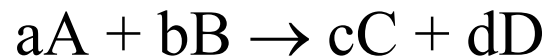
Theoretical maximum
at 1000 hPa, 293 K

(for $r_i, r_j = 0.2$ nm, $E_0 = 0$ and no steric hindrance): $k \approx 2 \times 10^{-10}$ cm³/molec/s

1.2 Homogeneous gas-phase reactions

Reactions can be unimolecular, bi- or termolecular.

The rate law of a reaction of the general form



Is defined as

$$\text{Rate (dt.: Rate)} = -dc_A/dt/a = -dc_B/dt/b = dc_C/dt/c = dc_D/dt/d$$

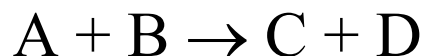
example:



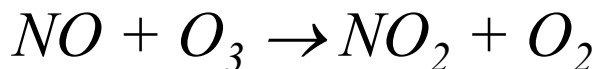
$$\text{Rate} = -dc_{NO}/dt/2 = -dc_{O_2}/dt = dc_{NO_2}/dt/2$$

Second order

Usually: bimolecular



example:



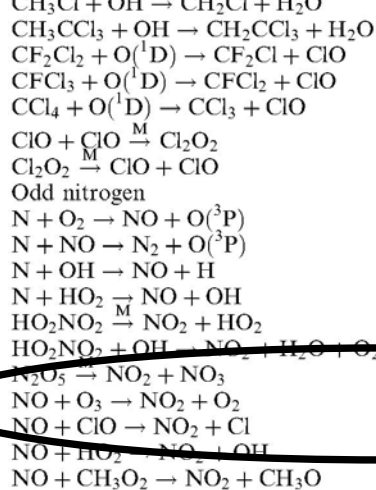
$$dc_C/dt = -dc_A/dt = -dc_B/dt = k^{(2)} c_A^1 c_B^1$$

Reaction rate: second order (1+1=2)

Reaction rate coefficient (denoting the order): $k^{(2)}$

The reaction order is given by the sum of the exponentials, $n+m+\dots$, of the concentration terms in the rate law of the form $-dc_A/dt = k c_A^n c_B^m$ ($n = \text{zero or integer or fraction}^*$)
It is determined empirically.

* ,overall‘ reactions only



$$\begin{aligned}
R_{\text{Cl14}} &= 2.1 \times 10^{-12} \exp(-1150/T) \\
R_{\text{Cl15}} &= 1.8 \times 10^{-12} \exp(-1550/T) \\
R_{\text{Cl16}} &= 1.4 \times 10^{-10} \\
R_{\text{Cl17}} &= 2.3 \times 10^{-10} \\
R_{\text{Cl18}} &= 3.3 \times 10^{-10}
\end{aligned}$$

$$\begin{aligned}
R_{\text{Cl19}} &= f(K_0, K_\infty), K_0 = 1.9 \times 10^{-32} [M] (300/T)^{3.9}, K_\infty = 7.0 \times 10^{-32} \\
R_{\text{Cl20}} &= R_{\text{Cl19}} / (3.0 \times 10^{-27} \exp(8450/T))
\end{aligned}$$

$$R_{\text{N1}} = 1.5 \times 10^{-11} \exp(-3600/T)$$

$$R_{\text{N2}} = 3.4 \times 10^{-11}$$

$$R_{\text{N3}} = 5.3 \times 10^{-11}$$

$$R_{\text{N4}} = 2.0 \times 10^{-10}$$

$$R_{\text{N5}} = R_{\text{N16}} / (2.1 \times 10^{-27} \exp(10900/T))$$

$$R_{\text{N6}} = 1.2 \times 10^{-12} \exp(380/T)$$

$$R_{\text{N7}} = R_{\text{N18}} / (4.0 \times 10^{-27} \exp(10930/T))$$

$$R_{\text{N8}} = 2.0 \times 10^{-12} \exp(-1400/T)$$

$$R_{\text{N9}} = 6.4 \times 10^{-12} \exp(290/T)$$

$$R_{\text{N10}} = 3.7 \times 10^{-12} \exp(250/T)$$

$$R_{\text{N11}} = 4.2 \times 10^{-12} \exp(180/T)$$

k_T

Example



$$k = A e^{[-(E_a/R)T]}$$

Arrhenius expression, preexponential factor A, activation energy E_a

$k(T^{-1}) \rightarrow$ slope $m = -E_a/R$, intercept $\ln A$, $E_a/R > 0 \leftrightarrow$ faster at higher T

$$A = 2 \times 10^{-12} \text{ cm}^3/\text{molec/s}$$

$$E/R = -1400 \text{ K}$$

\rightarrow

$$k_{298 \text{ K}} = 1.8 \times 10^{-14} \text{ cm}^3/\text{molec/s}$$

$$k_{230 \text{ K}} = 0.45 \times 10^{-14} \text{ cm}^3/\text{molec/s}$$

2. Exception:

First order (but not unimolecular):



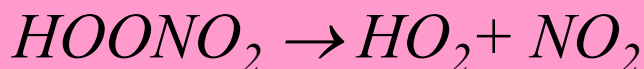
or if $c_B \gg c_A$



Here, M stands for any molecule or atom (i.e. N_2 , O_2 ,...), not transformed but required to absorb excess energy, e.g. of an activated intermediate state:



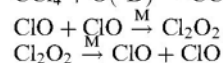
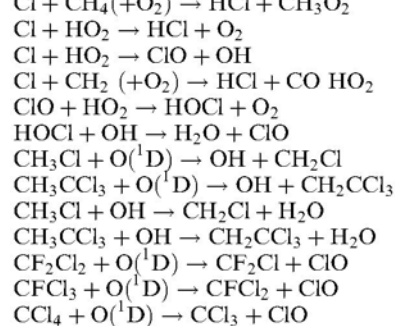
Example: thermic dissociation



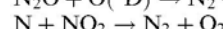
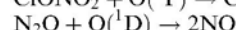
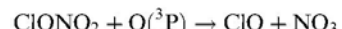
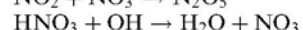
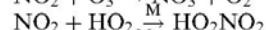
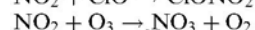
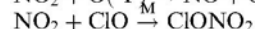
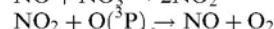
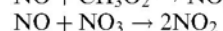
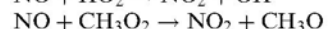
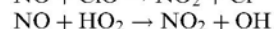
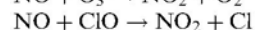
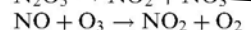
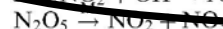
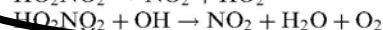
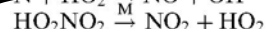
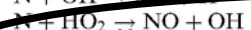
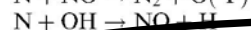
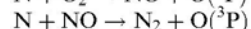
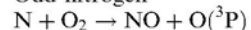
Reaction rate

$$dc_C/dt = -dc_A/dt = k^{(1)}c_A^1$$

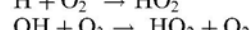
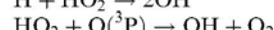
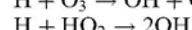
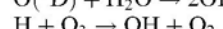
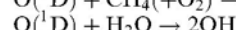
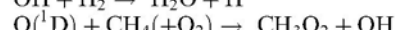
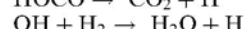
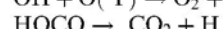
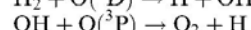
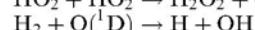
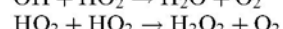
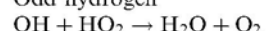
$$k^{(1)} = k^{(2)}c_M^{-1} \quad (c_M^1 \text{ constant})$$



Odd nitrogen



Odd hydrogen



$$\begin{aligned} R_{C6} &= 1.1 \times 10^{-11} \exp(-1400/T) \\ R_{C7} &= 1.8 \times 10^{-11} \exp(170/T) \\ R_{C8} &= 4.1 \times 10^{-11} \exp(-450/T) \\ R_{C9} &= 8.1 \times 10^{-11} \exp(-30/T) \\ R_{C10} &= 4.8 \times 10^{-13} \exp(700/T) \\ R_{C11} &= 3.0 \times 10^{-12} \exp(-500/T) \\ R_{C12} &= 4.0 \times 10^{-10} \\ R_{C13} &= 4.0 \times 10^{-10} \\ R_{C14} &= 2.1 \times 10^{-12} \exp(-1150/T) \\ R_{C15} &= 1.8 \times 10^{-12} \exp(-1550/T) \\ R_{C16} &= 1.4 \times 10^{-10} \\ R_{C17} &= 2.3 \times 10^{-10} \\ R_{C18} &= 3.3 \times 10^{-10} \end{aligned}$$

$$\begin{aligned} R_{C19} &= f(K_0, K_\infty), K_0 = 1.9 \times 10^{-32} [M] (300/T)^{3.9}, K_\infty \\ R_{C20} &= R_{C19} / (3.0 \times 10^{-27} \exp(8450/T)) \end{aligned}$$

$$R_{N1} = 1.5 \times 10^{-11} \exp(-3600/T)$$

$$R_{N2} = 3.4 \times 10^{-11}$$

$$R_{N3} = 5.3 \times 10^{-11}$$

$$R_{N4} = 2.0 \times 10^{-10}$$

$$R_{N5} = R_{N16} / (2.1 \times 10^{-27} \exp(10900/T))$$

$$R_{N6} = 1.3 \times 10^{-12} \exp(380/T)$$

$$R_{N7} = R_{N18} / (4.6 \times 10^{-13} \exp(10930/T))$$

$$R_{N8} = 2.0 \times 10^{-12} \exp(-1400/T)$$

$$R_{N9} = 6.4 \times 10^{-12} \exp(290/T)$$

$$R_{N10} = 3.7 \times 10^{-12} \exp(250/T)$$

$$R_{N11} = 4.2 \times 10^{-12} \exp(180/T)$$

$$R_{N12} = 1.5 \times 10^{-11} \exp(170/T)$$

$$R_{N13} = 6.5 \times 10^{-12} \exp(120/T)$$

$$R_{N14} = f(K_0, K_\infty), K_0 = 1.8 \times 10^{-31} [M] (300/T)^{3.4}, K_\infty$$

$$R_{N15} = 1.2 \times 10^{-13} \exp(-2450/T)$$

$$R_{N16} = f(K_0, K_\infty), K_0 = 1.8 \times 10^{-31} [M] (300/T)^{3.2}, K_\infty$$

$$R_{N17} = f(K_0, K_\infty), K_0 = 2.6 \times 10^{-30} [M] (300/T)^{3.2}, K_\infty$$

$$R_{N18} = f(K_0, K_\infty), K_0 = 2.2 \times 10^{-30} [M] (300/T)^{3.9}, K_\infty$$

$$R_{N19} = K_1 + K_2 / (1 + K_2/K_3), K_1 = 7.2 \times 10^{-15} \exp(7800/T), K_3 = 4.1 \times 10^{-16} \exp(1440/T)$$

$$R_{N20} = 2.9 \times 10^{-12} \exp(-800/T)$$

$$R_{N21} = 6.7 \times 10^{-11}$$

$$R_{N22} = 4.9 \times 10^{-11}$$

$$R_{N23} = 3.0 \times 10^{-12}$$

$$R_{H1} = 4.8 \times 10^{-11} \exp(250/T)$$

$$R_{H2} = 2.3 \times 10^{-13} \exp(600/T) + 1.7 \times 10^{-33} [M] \exp(10000/T)$$

$$R_{H3} = 1.0 \times 10^{-10}$$

$$R_{H4} = 2.2 \times 10^{-11} \exp(120/T)$$

$$R_{H5} = 1.5 \times 10^{-13} (1 + 0.6p/101400)$$

$$R_{H6} = 5.5 \times 10^{-12} \exp(-2000/T)$$

$$R_{H7} = 1.4 \times 10^{-10}$$

$$R_{H8} = 2.2 \times 10^{-10}$$

$$R_{H9} = 1.4 \times 10^{-10} \exp(-470/T)$$

$$R_{H10} = 7.1685 \times 10^{-11}$$

$$R_{H11} = 3.0 \times 10^{-11} \exp(200/T)$$

$$R_{H12} = 1.1 \times 10^{-14} \exp(-500/T)$$

$$R_{H13} = f(K_0, K_\infty), K_0 = 3.7 \times 10^{-32} [M] (300/T)^{1.6}, K_\infty$$

$$R_{H14} = 1.6 \times 10^{-12} \exp(-940/T)$$

Steil et al., 1998

Reaction orders



is 2nd order or – in case $dc_B/dt \approx 0$ - pseudo-1st order



is 2nd order or – in case $dc_B/dt \approx 0$ - pseudo-1st order or
– if $p \ll 1000$ hPa - between 2nd and 3rd order

The unit of a homogeneous gas-phase rate is [molec/cm³/s].

A 1st or pseudo-1st order rate law reads:

$$-dc_A/dt = dc_C/dt = k^{(1)} c_C, \text{ with } k^{(1)} [1/s].$$

A 2nd order rate law reads:

$$-dc_A/dt = dc_C/dt = k^{(2)} c_C c_D, \text{ with } k^{(2)} [\text{cm}^3/\text{molec}/s].$$

A 3rd order rate law reads e.g.:

$$-dc_A/dt = k^{(3)} c_C^2 c_D, \text{ with } k^{(3)} [\text{cm}^6/\text{molec}^2/s].$$

3. **between 2nd and 3rd order:** In the range of pressures and temperatures in the upper troposphere and stratosphere (c_M small) many reactions are in the region between 2nd and 3rd order reactions:

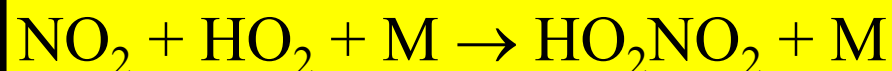
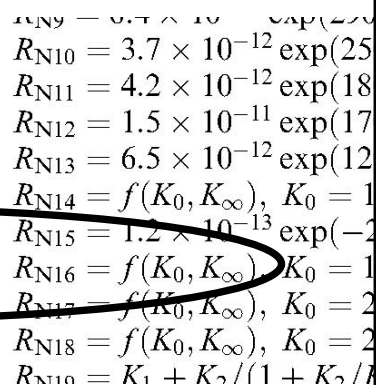
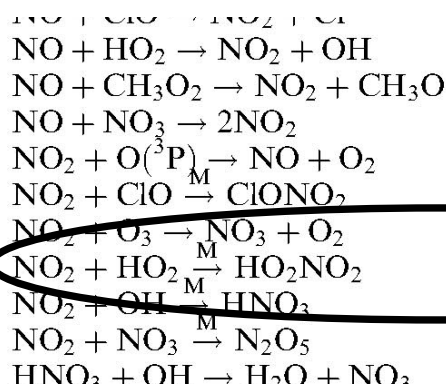
$$k = \frac{k_0 M}{1 + k_0 M / k_\infty} F_C^{(1 + \log(k_0 M / k_\infty))^2} \quad F_C \approx 0.6$$

$$k_0(T) = k_0^{300K} \left(\frac{T}{300K} \right)^{-n}, \quad k_\infty(T) = k_\infty^{300K} \left(\frac{T}{300K} \right)^{-m}$$

k_0, k_∞ = low, high pressure limiting rate constant

(Troe, 1979)

F_C = broadening factor of falloff curve between 2nd and 3rd order in the range of atmospheric pressures



$K_0 = k_0(T)c_M, n = 3.2$

$K_\infty = k_\infty(T), m = 1.4 \rightarrow$

$k_{288 \text{ K}/1000 \text{ hPa}} = 1.9 \times 10^{-12} \text{ cm}^3/\text{molec/s}$

$k_{288 \text{ K}/500 \text{ hPa}} = 1.5 \times 10^{-12} \text{ cm}^3/\text{molec/s}$

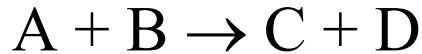
$k_{230 \text{ K}/500 \text{ hPa}} = 2.3 \times 10^{-12} \text{ cm}^3/\text{molec/s}$

$k_{230 \text{ K}/1000 \text{ hPa}} = 3.0 \times 10^{-12} \text{ cm}^3/\text{molec/s}$

1.3 Combinations of thermic reactions- Quasi steady state approximation

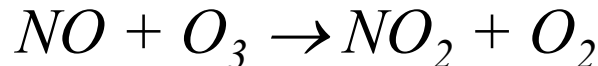
Second order

Usually: bimolecular



A, B, C, D molecules or radicals

example:



Reaction rate: second order (1+1=2)

$$dc_C/dt = -dc_A/dt = -dc_B/dt = k c_A^1 c_B^1$$

Reaction rate coefficient: $k^{(2)}$

$$k = A \exp [-(E_a/R)T]$$

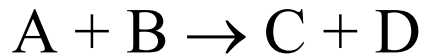
Arrhenius expression, preexponential factor A, activation energy E_a

$k(T^{-1}) \rightarrow$ slope $-E_a/R$, intercept $\ln A$, $E_a/R > 0 \leftrightarrow$ faster at higher T

1.3 Combinations of thermic reactions- Quasi steady state approximation

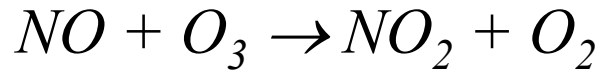
Second order

Usually: bimolecular



A, B, C, D molecules or radicals

example:



Reaction rate: second order (1+)

$$dc_C/dt = -dc_A/dt = -dc_B/dt = k c_A c_B$$

Reaction rate coefficient: $k^{(2)}$

$$k = A \exp [-(E_a/R)T]$$

Arrhenius expression, preexponential factor

$k(T^{-1}) \rightarrow$ slope $-E_a/R$, intercept $\ln A$,

Very frequent:



Intermediate C^* :

$$dc_{C^*}/dt = k_1 c_A c_B - k_{-1} c_{C^*} - k_2 c_{C^*} c_M$$

Lifetime of C^* short $\rightarrow dc_{C^*}/dt = 0$

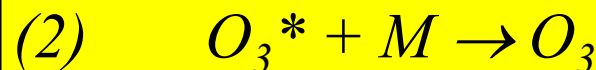
Hence:

$$k_1 c_A c_B = k_{-1} c_{C^*} + k_2 c_{C^*} c_M$$

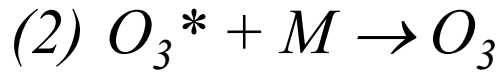
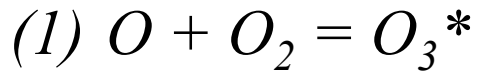
$$k_1 c_A c_B = c_{C^*} (k_{-1} + k_2 c_M)$$

$$c_{C^*} = k_1 c_A c_B / (k_{-1} + k_2 c_M)$$

Example



Example



$$c_{O_3^*} = k_1 c_O c_{O_2} / (k_{-1} + k_2 c_M)$$

$$dc_{O_3}/dt = k_2 c_{O_3^*} c_M$$

$$dc_{O_3}/dt = k_1 k_2 c_O c_{O_2} c_M / (k_{-1} + k_2 c_M) = \underbrace{[k_1 k_2 c_M / (k_{-1} + k_2 c_M)]}_{\mathbf{k}} c_O c_{O_2}$$

High pressure limit: $k_{-1} \approx 0 \rightarrow k_{\infty} = k_1$

Low pressure limit: $c_M \rightarrow 0 \rightarrow k_0 = k_1 k_2 / k_{-1}$

Reaction rate compilations

Compilations are considered the state-of-the-art description of atmospheric chemistry, because

- complete knowledge is summarized and assessed, and
- laboratory kineticists from around the world are involved

However,

- there is no guarantee for completeness, and
- some rate constants are very uncertain, but uncertainty is quantified

• Chemical kinetics and photochemical data for use in stratospheric modeling
Eval. No. 12 (JPL-Publ-97-4; NAS 1.26112606; NASA CR-97-112606), 1997
(updated 2005, <http://ntrs.nasa.gov/search.jsp>)

• Hydrocarbon reactions: Atkinson et al., Atmos. Chem. Phys. 6 (2006) 3625-4055

Chemical kinetics and photochemical data for use in stratospheric modeling
 Eval. No. 12 (JPL-Publ-97-4; NAS 1.26112606; NASA CR-97-112606), 1997
 (updated 2005)
<http://ntrs.nasa.gov/search.jsp>

$$k = A \exp (-E_a / RT)$$

Arrhenius expression

TABLE OF CONTENTS

INTRODUCTION	1
Basis of The Recommendations	3
Recent Changes and Current Needs of Laboratory Kinetics	3
Format of the Evaluation	3
Computer Access	3
Ox Reactions	3
Reactions of Singlet Oxygen	3
HO _x Reactions	4
NO _x Reactions	4
Oxidation of Organic Compounds	4
Halogen Reactions	5
SO _x Reactions	5
Metal Chemistry	5
Photochemical Data	5
Heterogeneous Chemistry	6
Gas Phase Enthalpy Data (Appendix 1)	7
Entropy Data (Appendix 2)	7
Solar Flux and Species Profiles (Appendix 3)	7
Data Formats	7
Bimolecular Reactions	7
Termolecular Reactions	8
Uncertainty Estimates	10
Units	11
References for the Introduction	13
RATE CONSTANT DATA	14
Table of Data for Second Order Reactions (Table 1)	14
Notes to Table 1	39
References for Table 1	101
Table of Data for Association Reactions (Table 2)	123
Notes to Table 2	127
References for Table 2	135
EQUILIBRIUM CONSTANTS	141
Format	141
Definitions	141
Notes to Table 3	143
References for Table 3	144
PHOTOCHEMICAL DATA	146
Discussion of Format and Error Estimates	146
O ₂ + hν → O + O	149
O ₃ + hν → O + O ₂	150
HO ₂ + hν → OH + H	152
H ₂ O + hν → H + OH	153
H ₂ O ₂ + hν → OH + OH	154
NO ₂ + hν → NO + O	155
NO ₃ + hν → NO + O ₂ (Φ ₁)	158
NO ₃ + hν → NO ₂ + O (Φ ₂)	158
N ₂ O + hν → N ₂ + O(¹ D)	160
N ₂ O ₅ + hν → Products	162
HONO + hν → OH + NO	163
HNO ₃ + hν → products	163
HO ₂ NO ₂ + hν → Products	165
CH ₂ O + hν → H + HCO (Φ ₁)	165

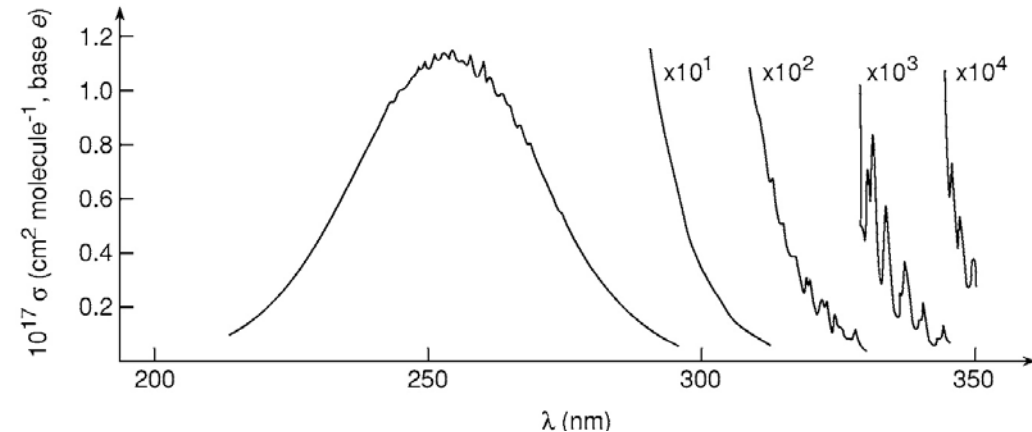
Table 1. (Continued)

Reaction	A-Factor ^a	E/R±(ΔE/R)	k(298 K) ^a	f(298) ^b	Notes
OH + HONO → H ₂ O + NO ₂	1.8x10 ⁻¹¹	390± ²⁰⁰ ₅₀₀	4.5x10 ⁻¹²	1.5	C 8
OH + HNO ₃ → H ₂ O + NO ₃	(See Note)			1.3	C 9
OH + HO ₂ NO ₂ → products	1.3x10 ⁻¹²	-(380± ²⁷⁰ ₅₀₀)	4.6x10 ⁻¹²	1.5	C10
OH + NH ₃ → H ₂ O + NH ₂	1.7x10 ⁻¹²	710±200	1.6x10 ⁻¹³	1.2	C11
HO ₂ + NO → NO ₂ + OH	3.5x10 ⁻¹²	-(250±50)	8.1x10 ⁻¹²	1.15	C12
HO ₂ + NO ₂ \xrightarrow{M} HO ₂ NO ₂	(See Table 2)				
HO ₂ + NO ₂ → HONO + O ₂	(See Note)				C13
HO ₂ + NO ₃ → products			3.5x10 ⁻¹²	1.5	C14
HO ₂ + NH ₂ → products			3.4x10 ⁻¹¹	2.0	C15
N + O ₂ → NO + O	1.5x10 ⁻¹¹	3600±400	8.5x10 ⁻¹⁷	1.25	C16
N + O ₃ → NO + O ₂			<2.0x10 ⁻¹⁶		C17
N + NO → N ₂ + O	2.1x10 ⁻¹¹	-(100±100)	3.0x10 ⁻¹¹	1.3	C18
N + NO ₂ → N ₂ O + O	5.8x10 ⁻¹²	-(220±100)	1.2x10 ⁻¹¹	1.5	C19
NO + O ₃ → NO ₂ + O ₂	2.0x10 ⁻¹²	1400±200	1.8x10 ⁻¹⁴	1.1	C20
NO + NO ₃ → 2NO ₂	1.5x10 ⁻¹¹	-(170±100)	2.6x10 ⁻¹¹	1.3	C21
NO ₂ + O ₃ → NO ₃ + O ₂	1.2x10 ⁻¹³	2450±150	3.2x10 ⁻¹⁷	1.15	C22
NO ₂ + NO ₃ → NO + NO ₂ + O ₂	(See Note)				C23
NO ₂ + NO ₃ \xrightarrow{M} N ₂ O ₅	(See Table 2)				
NO ₃ + NO ₃ → 2NO ₂ + O ₂	8.5x10 ⁻¹³	2450±500	2.3x10 ⁻¹⁶	1.5	C24
NH ₂ + O ₂ → products			<6.0x10 ⁻²¹		C25
NH ₂ + O ₃ → products	4.3x10 ⁻¹²	930±500	1.9x10 ⁻¹³	3.0	C26
NH ₂ + NO → products	4.0x10 ⁻¹²	-(450±150)	1.8x10 ⁻¹¹	1.3	C27
NH ₂ + NO ₂ → products	2.1x10 ⁻¹²	-(650±250)	1.9x10 ⁻¹¹	3.0	C28
NH + NO → products	4.9x10 ⁻¹¹	0±300	4.9x10 ⁻¹¹	1.5	C29
NH + NO ₂ → products	3.5x10 ⁻¹³	-(1140±500)	1.6x10 ⁻¹¹	2.0	C30

1.4 Photochemical reactions

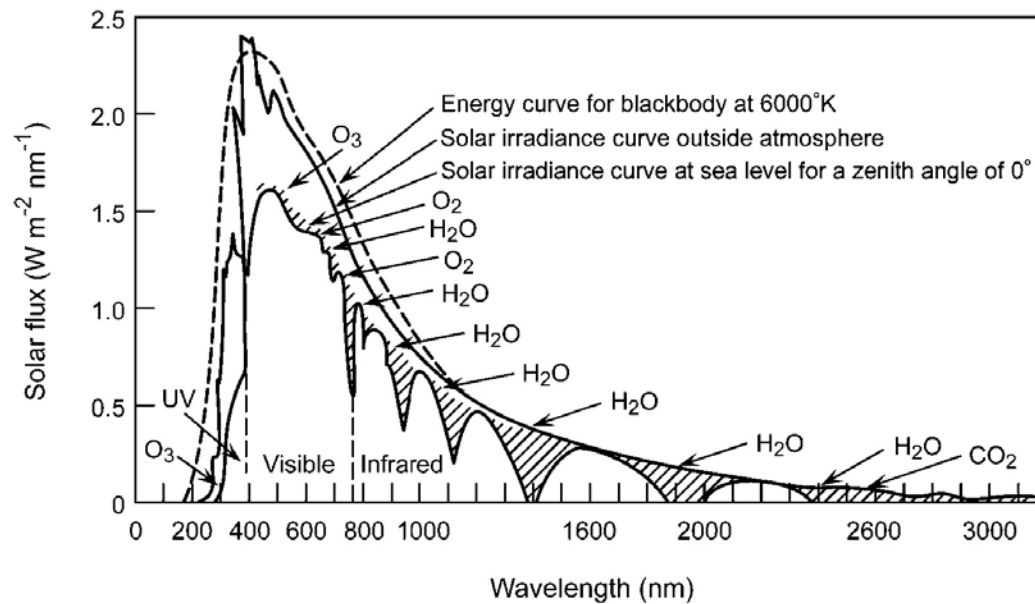
Radiation absorption in the atmosphere

Gaseous molecules absorb ultraviolet, visible and infrared light: O_3



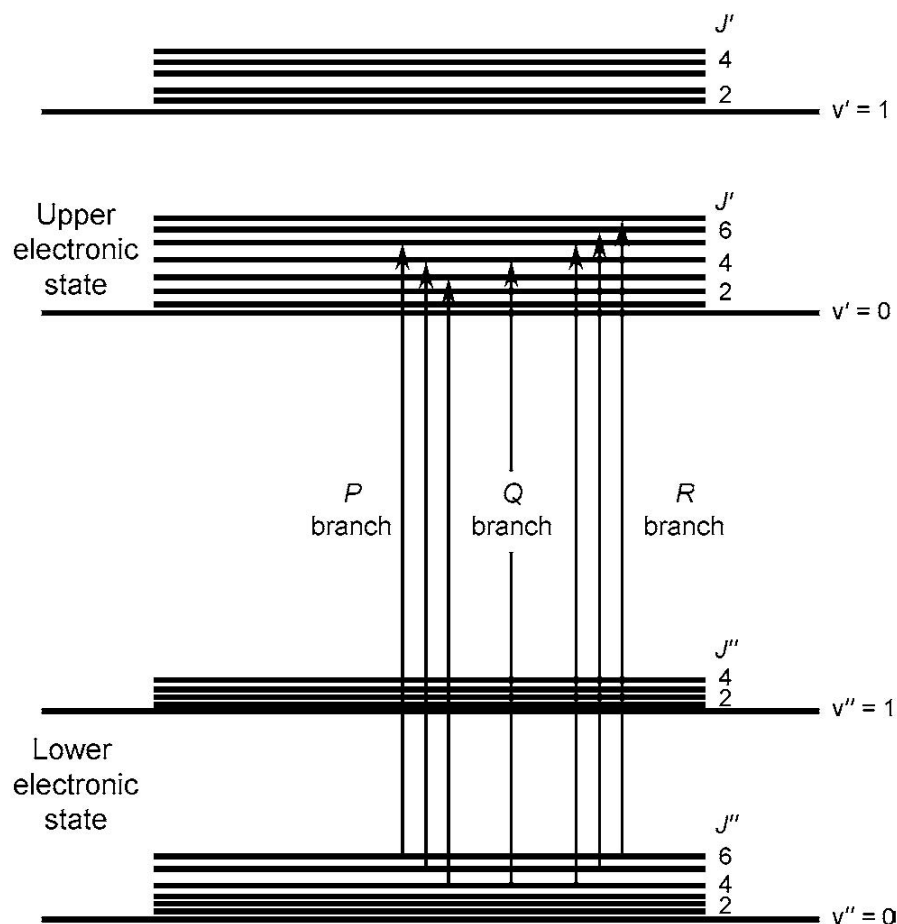
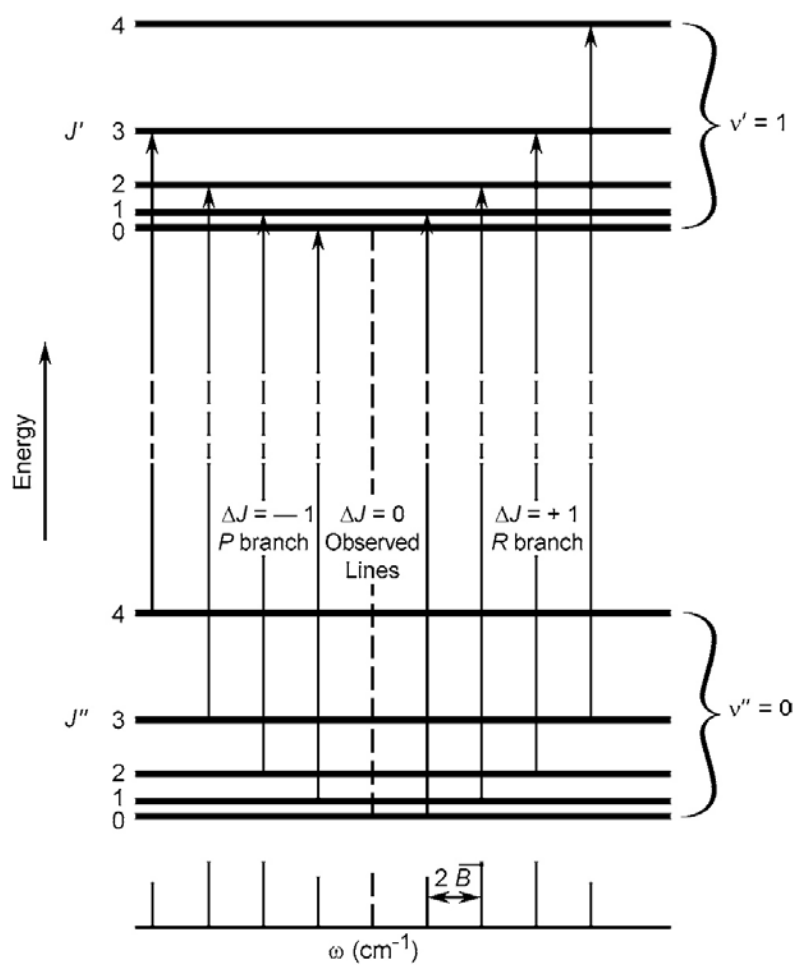
Consequences:

- photophysical and photochemical molecular processes
- change of spectrum:



Specific because of vibrational-rotational and electronic transitions:

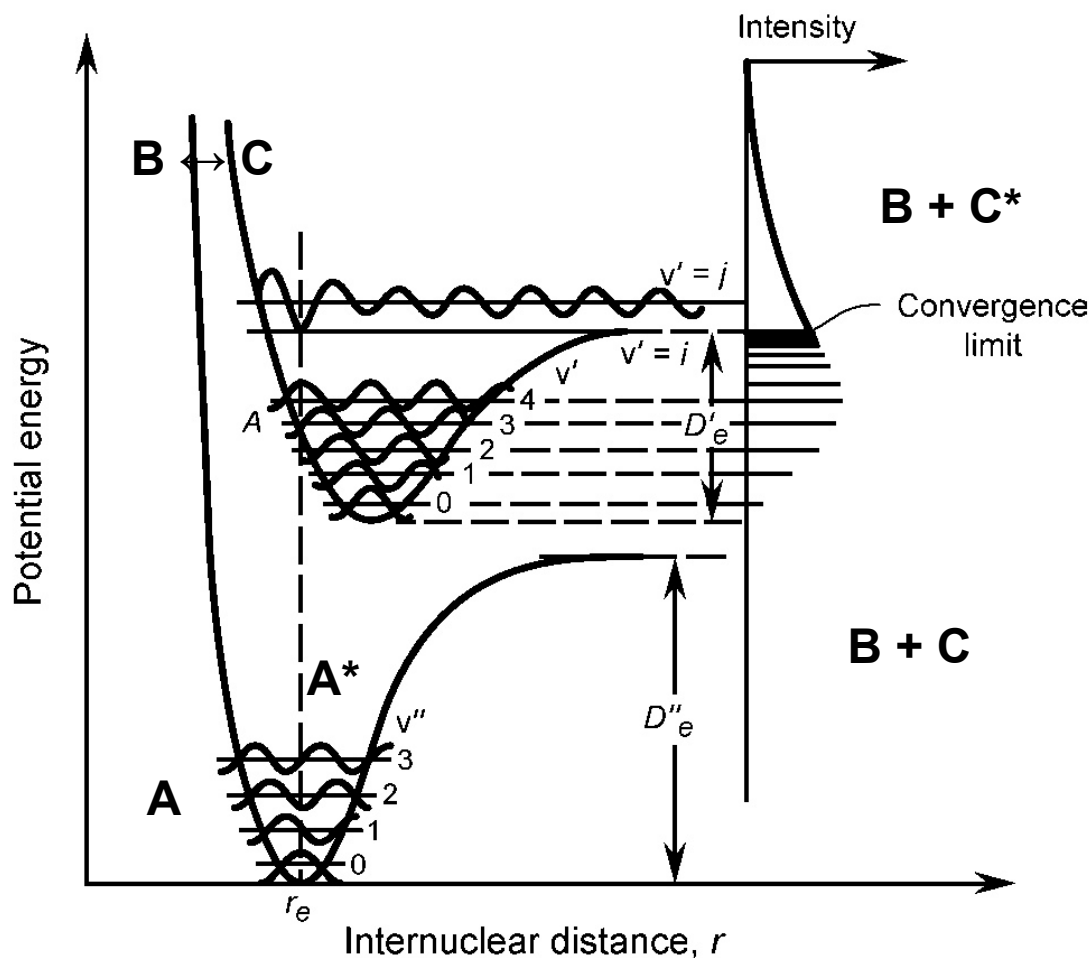
A series of vibrational-rotational states are populated (temperature dependent, Boltzmann) within electronic states. Vibrational-rotational (P, Q, R branches of change of J) and electronic transitions follow quantum rules (selection rules, 'allowed' or not allowed). These correspond to characteristic spectroscopic patterns. Therefore, emission and absorption spectroscopy allows for unambiguous identification of molecules.



Electronic excitation:



The energy barrier will be smaller and the distance between atoms will be larger for the excited molecule.

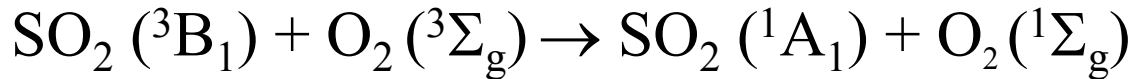


Fates of electronically excited molecules:

Photophysical primary processes: emission (fluorescence, phosphorescence), nonradiative energy transition into heat

Photochemical primary processes: dissociation, isomerization, re-arrangement, reaction

Example for photophysical process: $A^ + B \rightarrow A + B^*$:*



Example for photochemical process:



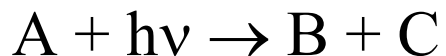
Def.:

Efficiency of a primary photochemical process i is given by its quantum yield ϕ_i := number of excited molecules by process i divided by total number of photons absorbed.

The sum of the ϕ_i of all the photochemical primary processes = 1.

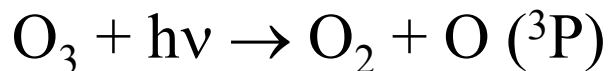
Photodissociations

Unimolecular



A, B, C, D molecules or radicals

example:



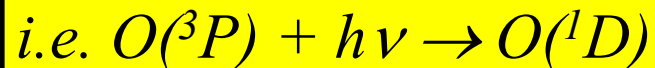
Reaction rate coefficient j (photolysis rate):

$$dc_C/dt = -dc_A/dt = j c_A$$

Ground state, A
/excitation state, A*:



Example:



Energy ranges, correspondence between energy and wavelength

$$\lambda = c/\nu$$

with frequency ν

$$\Delta E = hc/\lambda = hc\omega$$

Planck relationship (wavelength λ , wavenumber ω ,
Planck's constant $h = 6.626 \times 10^{-34}$ Js)

Commonly used energy units:

$$(\text{kJ mol}^{-1})$$

$$\times 0.2390 = \text{kcal mol}^{-1}$$

$$\times 0.0104 = \text{eV}$$

$$\times 83.59 = \text{cm}^{-1}$$

$$(\text{kcal mol}^{-1})$$

$$\times 4.184 = \text{kJ mol}^{-1}$$

$$\times 0.04336 = \text{eV}$$

$$\times 349.8 = \text{cm}^{-1}$$

$$(\text{cm}^{-1})$$

$$\times 1.196 \times 10^{-2} = \text{kJ mol}^{-1}$$

$$\times 2.859 \times 10^{-3} = \text{kcal mol}^{-1}$$

$$\times 1.240 \times 10^{-4} = \text{eV}$$

$$(\text{eV})$$

$$\times 96.49 = \text{kJ mol}^{-1}$$

$$\times 23.06 = \text{kcal mol}^{-1}$$

$$\times 8.064 \times 10^3 = \text{cm}^{-1}$$

Energy ranges, correspondence between energy and wavelength

$$\lambda = c/\nu$$

$$\Delta E = hc/\lambda = hc\omega$$

Planck relationship (wavelength λ , wavenumber ω , Planck's constant $h = 6.626 \times 10^{-34}$ Js)

Name	Typical wavelength or range of wavelengths (nm)	Typical range of frequencies ν (s^{-1})	Typical range of wavenumbers ω (cm^{-1})	Typical range of energies (kJ einstein $^{-1}$) ^a
Radiowave	$\sim 10^8$ – 10^{13}	$\sim 3 \times 10^4$ – 3×10^9	10^{-6} – 0.1	$\sim 10^{-3}$ – 10^{-8}
Microwave	$\sim 10^7$ – 10^8	$\sim 3 \times 10^9$ – 3×10^{10}	0.1 – 1	$\sim 10^{-2}$ – 10^{-3}
Far-infrared	$\sim 10^5$ – 10^7	$\sim 3 \times 10^{10}$ – 3×10^{12}	1 – 100	$\sim 10^{-2}$ – 1
Near-infrared	$\sim 10^3$ – 10^5	$\sim 3 \times 10^{12}$ – 3×10^{14}	10^2 – 10^4	~ 1 – 10^2
Visible				
Red	700	4.3×10^{14}	1.4×10^4	1.7×10^2
Orange	620	4.8×10^{14}	1.6×10^4	1.9×10^2
Yellow	580	5.2×10^{14}	1.7×10^4	2.1×10^2
Green	530	5.7×10^{14}	1.9×10^4	2.3×10^2
Blue	470	6.4×10^{14}	2.1×10^4	2.5×10^2
Violet	420	7.1×10^{14}	2.4×10^4	2.8×10^2
Near-ultraviolet	400–200	$(7.5$ – $15.0) \times 10^{14}$	$(2.5$ – $5) \times 10^4$	$(3.0$ – $6.0) \times 10^2$
Vacuum ultraviolet	~ 200 – 50	$(1.5$ – $6.0) \times 10^{15}$	$(5$ – $20) \times 10^4$	$\sim (6.0$ – $24) \times 10^2$
X-Ray	~ 50 – 0.1	$\sim (0.6$ – $300) \times 10^{16}$	$(0.2$ – $100) \times 10^6$	$\sim 10^3$ – 10^6
γ -Ray	≤ 0.1	$\sim 3 \times 10^{18}$	$\geq 10^8$	$> 10^6$

^a For kcal einstein $^{-1}$, divide by 4.184 (1 cal = 4.184 J).

The rate of photochemical reactions

Absorption

$$\ln(I_0/I) = \sigma Nd$$

Beer-Lambert law

$$I/I_0 = e^{(-\sigma Nd)}$$

absorption cross section σ (cm², default: base e)

molecule concentration N (cm⁻³), depth of absorptive layer d (cm)

optical depth $OD = \sigma Nd$

Caution: Most measurements are made to the

base 10 ($\log(I_0/I) = \sigma_{10} Nd$) $\rightarrow \times 2.303$ to reach base e

The photolysis rate

The photolysis rate, j (s^{-1}), is given by:

$$j = \int_{\lambda} \phi(\lambda) \sigma(\lambda) L(\lambda) d\lambda$$

- quantum yield $\phi(\lambda)$ (),
- absorption cross section σ (cm^2),
- actinic flux, $L(\lambda)$ ($\text{cm}^{-2} \text{s}^{-1}$)

L is the total intensity of effective light (direct + scattered + reflected, spherically integrated).

The photolysis rate

$L(\lambda)$ is a function of the solar zenith angle, cloudiness, aerosol concentration, and surface albedo*.

* Type of surface	Albedo
Snow	0.69 0.93 ^b 0.9–1.0
Ocean	0.07 ^b 0.06–0.08 ^a
Forests	0.06–0.18 ^b 0.02 ^a 0.17
Fields and meadows	0.03–0.04 ^a
Desert	0.06–0.09 ^a
Salt flats	0.57–0.65 ^a

^a Minimum reflectivities at 370 nm.

^b Measured with respect to NO₂ ph

$$\frac{dn_A}{dt} = -n_A \underbrace{\int_{\lambda} \phi(\lambda) \sigma(\lambda) \left(\int_0^{2\pi} \int_0^{\pi} L(\lambda, \theta, \varphi) \sin \theta d\theta d\varphi \right) d\lambda}_{j_A}$$

Typical j-values for midlatitude noontime equinox conditions range from $\sim 1 \cdot 10^{-5} \text{ s}^{-1}$ for $j_{\text{O}(1\text{D})}$ to $\sim 0.2 \text{ s}^{-1}$ for j_{NO_3} . The actinic flux under these conditions is about $2 \cdot 10^{14} \text{ photons cm}^{-2} \text{ s}^{-1}$ at 315-320 nm, and $7 \cdot 10^{14} \text{ photons cm}^{-2} \text{ s}^{-1}$ at 360-365 nm.

The photolysis rate

L is measured using a (2π) radiometer or by measuring the photolytic decay (so-called chemical actinometry).

Its value can be estimated via tabulated values of ϕ and σ for intervals of λ and estimates of $L(\lambda)$ for given conditions.

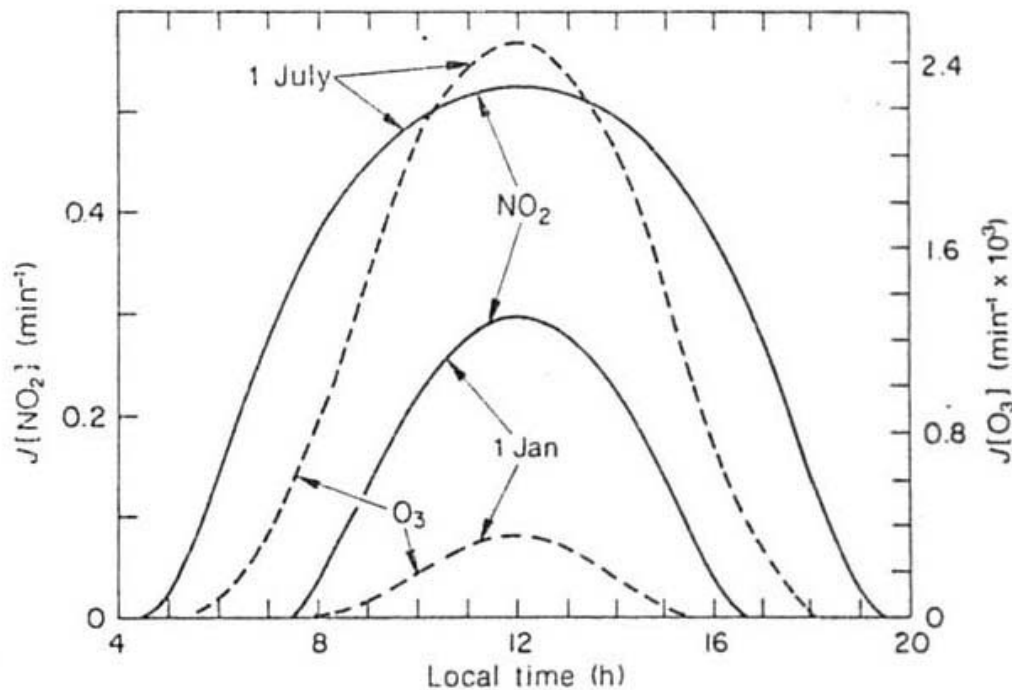


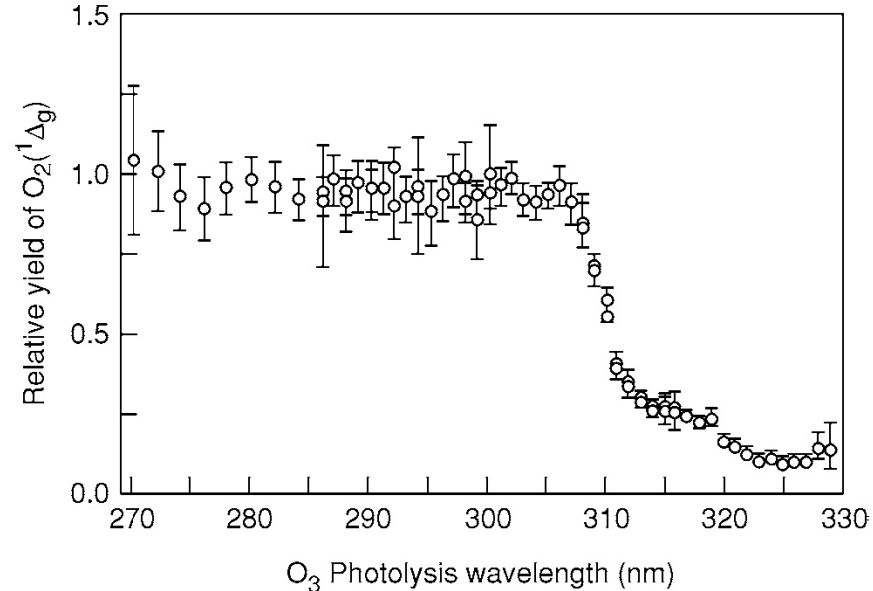
Fig. 3 Theoretical diurnal variation of the $J[\text{NO}_2]$ and $J[\text{O}_3]$ values for midsummer and midwinter at 40°N latitude near sea level. (Calvert, 1985)

Example j_{O_3} : Quantum yield $\phi_i(\lambda)$ for $i = O_3 + h\nu \rightarrow O(^1D) + O_2(^3\Sigma_g^-)$

TABLE 4.6 Parameterization of Quantum Yields for $O(^1D)$ Production from O_3 Photolysis in the 306- to 329-nm Region at Various Temperatures^a

Wavelength (nm)	<i>A</i>	<i>B</i>
306	0.80	9.84
307	0.78	1.44
308	0.87	53.1
309	0.76	73.9
310	1.31	305.5
311	2.37	600
312	5.8	925.9
313	11.4	1191
314	20.1	1423
315	26.4	1514
316	26.8	1512
317	26.8	1542
318	28.33	1604
319	30.6	1604
320	44.4	1866
321	50.2	1931
322	27.8	1882
323	74.1	2329
324	868	3085
325	0.37	689
326	0.24	619
327	0.068	258
328	26.16	2131
329	0.15	470

^a Using the quantum yield expression recommended by Talukdar *et al.*, 1998: $\phi = 0.06 + Ae^{-B/T}$.



$$\phi_{O_3 \rightarrow O(^1D)}(310-320 \text{ nm}) \approx 0.2$$

Example j_{O_3} : Absorption cross section $\sigma(\lambda, T)$ of O_3

TABLE 4.3 Ozone Absorption Cross Sections (Base e)^a

Wavelength (nm)	10 ²⁰ σ (cm ² molecule ⁻¹)		Wavelength (nm)
	$T = 226$ K	$T = 298$ K	
185.0	64.37	65.37	262.0
186.0	62.59	61.87	263.0
187.0	59.33	59.41	264.0
188.0	56.55	56.59	265.0
189.0	54.63	54.24	266.0
190.0	51.63	51.14	267.0
191.0	48.42	48.80	268.0
192.0	45.95	46.06	269.0
193.0	43.12	43.36	270.0
194.0	40.88	40.66	271.0
195.0	38.27	38.64	272.0
196.0	36.42	36.73	273.0
197.0	34.63	35.00	274.0
198.0	33.33	33.49	275.0
199.0	32.13	32.09	276.0
200.0	31.45	31.54	277.0
201.0	31.26	31.15	278.0
202.0	31.56	31.79	279.0
203.0	32.55	32.51	280.0
204.0	34.00	33.65	281.0
205.0	36.23	35.85	282.0
206.0	38.87	38.55	283.0
207.0	42.39	42.00	284.0
208.0	46.84	46.40	285.0
209.0	51.88	51.18	286.0
210.0	58.06	57.16	287.0
211.0	65.28	64.02	288.0
212.0	73.12	71.94	289.0
213.0	82.58	81.04	290.0
214.0	92.55	90.96	291.0
215.0	104.1	102.3	292.0
216.0	116.9	114.6	293.0
217.0	131.4	128.7	294.0
218.0	146.4	143.9	295.0
219.0	163.8	160.1	296.0
220.0	179.9	178.5	297.0
221.0	200.0	198.2	298.0
222.0	221.7	220.0	299.0
223.0	244.3	242.9	300.0
224.0	268.8	268.4	301.0
225.0	296.3	294.3	302.0
226.0	323.9	322.6	303.0
227.0	354.2	351.3	304.0
228.0	385.7	382.9	305.0
229.0	416.4	414.1	306.0
230.0	450.6	447.6	307.0
231.0	485.9	481.4	308.0
232.0	523.0	518.1	309.0

TABLE 4.4 Ozone Absorption Cross Sections^{a, b} as a Function of Temperature Averaged over the Spectral Intervals Shown

Wavelength range (nm)	Parameters		
	a	b	c
277.778–281.690	4.0293×10^2	$+4.3819 \times 10^{-2}$	0
281.690–285.714	2.7776×10^2	$+6.3125 \times 10^{-2}$	0
285.714–289.855	1.8417×10^2	-9.6665×10^{-2}	2.1026×10^{-4}
289.855–294.118	1.1300×10^2	-1.0700×10^{-1}	3.2697×10^{-4}
294.118–298.507	6.5087×10	-8.0018×10^{-2}	2.2679×10^{-4}
298.507–303.030	3.6161×10	-6.7156×10^{-2}	3.3314×10^{-4}
303.030–307.692	1.9615×10	-4.4193×10^{-2}	2.0338×10^{-4}
307.692–312.5	1.0459×10	-2.8831×10^{-2}	1.3909×10^{-4}
312.5–317.5	5.4715	-2.0092×10^{-2}	9.8870×10^{-5}
317.5–322.5	2.7569	-1.0067×10^{-2}	2.9515×10^{-5}
322.5–327.5	1.3527	-5.7513×10^{-3}	1.1088×10^{-5}
327.5–332.5	6.9373×10^{-1}	-2.9792×10^{-3}	3.1038×10^{-6}
332.5–337.5	3.2091×10^{-1}	-1.9502×10^{-3}	5.6456×10^{-6}
337.5–342.5	1.4484×10^{-1}	-1.1025×10^{-3}	2.8818×10^{-6}
342.5–347.5	7.5780×10^{-2}	-5.7359×10^{-4}	1.6055×10^{-6}

^a $\sigma(O_3, T) = a + b(T - 230) + C(T - 230)^2$; T is in K; $\sigma(O_3)$ is in units of 10^{-20} cm² molecule⁻¹ (base e).

^b From Molina and Molina (1986).

Example:

$\sigma_{O_3}(310\text{-}320 \text{ nm}) \approx 60 \times 10^{-20} \text{ cm}^2/\text{molecule}$ for $T = 298 \text{ K}$

Data src.: Finlayson-Pitts & Pitts, 1998

Radiation absorption in the atmosphere

Actinic flux $L(\lambda)$ - Example:

For $z = 15$ km and solar zenith angle of 40° :

$$L(310-320 \text{ nm}) = (1.69 + 2.08 + 2.35 + 2.88 + 2.95) \times 10^{14} \text{ v cm}^{-2} \text{ s}^{-1}$$

Order of magnitude estimate of $j_{\text{O}_3 \rightarrow \text{O}(^1\text{D})}$ for a selected wavelength interval:

$$j_{\text{O}_3 \rightarrow \text{O}(^1\text{D})}(310-320 \text{ nm}) \approx 0.2 \times 60 \times 10^{-20} \times 12 \times 10^{14} \text{ s}^{-1} \approx 10^{-4} \text{ s}^{-1}$$

